

## Solution I

- 1) When comparing two satellites of the same object the ratio form of Kepler's Third Law (K3, periods) is good.

$$\left(\frac{P_1}{P_2}\right)^2 = \left(\frac{a_1}{a_2}\right)^3 \quad \text{or} \quad \left(\frac{T_1}{T_2}\right)^2 = \left(\frac{r_1}{r_2}\right)^3$$

In this form, you can use any units for  $P_1$  &  $P_2$  (period or time of one orbit) as long as both are the same, and any units for  $a_1$  &  $a_2$  (semi-major axis, orbital radius, or avg. distance).

$$P_1 = \frac{27.32 \text{ days} \left| \frac{24 \text{ hrs}}{1 \text{ day}} \right| \frac{60 \text{ min}}{1 \text{ hr}}}{\times 10^4 \text{ min}} = 3.93408$$

$$P_2 = 92.69 \text{ min}$$

$$a_1 = D_{EM} = 3.844 \times 10^8 \text{ m}$$

$$a_2 = ?$$

$$\text{Note: } h = a_2 - R_E$$

$$\left(\frac{P_1}{P_2}\right)^2 = \left(\frac{a_1}{a_2}\right)^3 \rightarrow a_2 = a_1 \left(\frac{P_2}{P_1}\right)^{2/3} = 3.844E8 \left(\frac{92.69}{3.93408E4}\right)^{2/3}$$

$$a_2 = 6.8062E6 \rightarrow h = 6.8062E6 - 6.371E6 = \boxed{4.35E5 \text{ m}}$$

## Solution II

Look at ISS's orbital velocity

$$V_{\text{orbit}} = \sqrt{\frac{GM}{r}}$$

$$\frac{2\pi r}{t} = \sqrt{\frac{GM_E}{r}}$$

$$\frac{4\pi^2 r^2}{t^2} = \frac{GM_E}{r}$$

$$r^3 = \frac{GM_E}{4\pi^2} t^2$$

$$r = \left( \frac{GM_E}{4\pi^2} t^2 \right)^{1/3}$$

$$r = 6.78 \times 10^6 \text{ m}$$

$$h = r - R_E = (6.78 - 6.371) \times 10^6 \text{ m}$$

$$= \boxed{4.09 \times 10^5 \text{ m}}$$

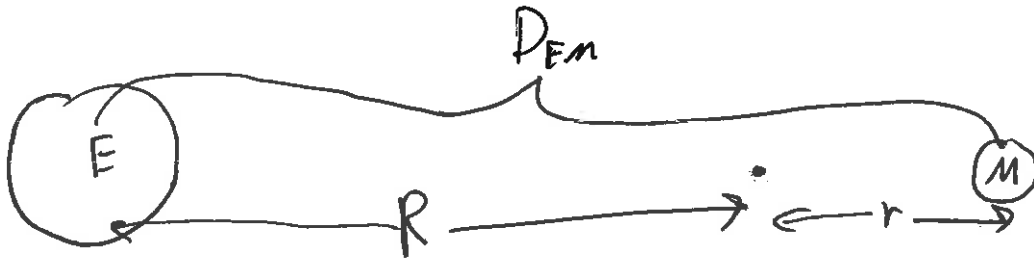
$$G = 6.67 \times 10^{-11}$$

$$M_E = 5.972 \times 10^{24} \text{ kg}$$

$$t = \frac{92.69 \text{ min}}{1 \text{ min}} \times 60 \text{ s} = 5,561.4 \text{ s}$$

difference of <10% from  
other answer is reasonable,  
probably from rounding.

2)



I'm going to guess in this picture that the point is closer to the Moon. Earth to the point is  $R$ , Moon to the point is  $r (= D_{EM} - R)$ . At this point, the gravities cancel out, so:

$$F_{gE} = F_{gM}$$

$$\frac{G M_E m}{R^2} = \frac{G M_m m}{r^2}$$

$$\frac{M_E}{R^2} = \frac{M_m}{(D_{EM} - R)^2}$$

you can now do algebra to find a quadratic equation, or use solver or Wolfram Alpha.

$$\frac{M_E}{R^2} - \frac{M_m}{(D_{EM} - R)^2} = 0$$

$$M_E = 5.972 \times 10^{24} \text{ kg} \quad M_m = 7.348 \times 10^{22} \text{ kg}$$

$$D_{EM} = 3.844 \times 10^8 \text{ m}$$

solver...

$$R = 3.460 \times 10^8 \text{ m}$$

Sanity check:  
Is it closer to  
E or M?

Or if you want an algebraic solution (and it seems like the NSpine calculator solver struggles with this),...

$$\frac{M_E}{R^2} = \frac{M_M}{(D_{EM} - R)^2}$$

$$\frac{(D_{EM} - R)^2}{R^2} = \frac{M_M}{M_E}$$

$$\left(\frac{D_{EM} - R}{R}\right)^2 = \frac{M_M}{M_E}$$

$$\left(\frac{D_{EM}}{R} - \frac{R}{R}\right)^2 = \frac{M_M}{M_E}$$

$$\left(\frac{D_{EM}}{R} - 1\right)^2 = \frac{M_M}{M_E}$$

$$\frac{D_{EM}}{R} - 1 = \sqrt{\frac{M_M}{M_E}}$$

$$\frac{D_{EM}}{R} = \sqrt{\frac{M_M}{M_E}} + 1$$

$$\frac{D_{EM}}{R} = \frac{\sqrt{M_M}}{\sqrt{M_E}} + \frac{\sqrt{M_E}}{\sqrt{M_E}}$$

$$\frac{D_{EM}}{R} = \frac{\sqrt{M_M} + \sqrt{M_E}}{\sqrt{M_E}}$$

$$\frac{R}{D_{EM}} = \frac{\sqrt{M_E}}{\sqrt{M_M} + \sqrt{M_E}}$$

$$R = D_{EM} \left( \frac{\sqrt{M_E}}{\sqrt{M_M} + \sqrt{M_E}} \right)$$

$$3) \text{ a) } V_{esc} = \sqrt{\frac{2GM}{r}}$$

$$G = 6.67 \times 10^{-11}, \quad M_E = 5.972 \times 10^{24}, \quad R_E = 6.371 \times 10^6 \text{ m}$$

$$V_{esc} = 1.118 \times 10^4 \text{ m/s} \approx 11 \text{ km/s}$$

$$\text{b) } V_{esc} = \sqrt{\frac{2GM}{r}}$$

$$M_S = 1.989 \times 10^{30} \text{ kg}, \quad R = D_{ES} = 1.496 \times 10^8 \text{ m}$$

$$V_{esc} = 4.211 \times 10^4 \text{ m/s}$$

Or could be asking how much speed added to Earth's orbital speed

$$V_{esc} = V + V_{orbit}$$

$$V = V_{esc} - V_{orbit}$$

$$= \sqrt{\frac{2GM}{r}} - \sqrt{\frac{GM}{r}}$$

$$= \sqrt{\frac{GM}{r}} = 2.978 \times 10^4 \text{ m/s}$$